

# Uncovering the hidden: Leveraging sub-pixel spectral diversity to estimate plant diversity from space Christian Rossi<sup>1,2</sup> & Hamed Gholizadeh<sup>1</sup>

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### **Background:**

- Spectral diversity has emerged as a valuable proxy for plant diversity.
- Pixel size of spaceborne data limits the estimation of local plant • diversity via spectral diversity in grasslands.

#### **Assumption:**

Spectral signature of a pixel is a linear combination of spectra of unique spectral species present within that pixel (i.e., plant endmembers  $s_1$ ,  $s_2$ ,  $s_3$ ) weighted by their corresponding abundances ( $\alpha_1$ ,  $\alpha_2$ ,  $\alpha_3$ ).



### Method:

- Applying spectral unmixing to derive the abundance ( $\alpha_i$ ) of unique spectral species (s<sub>i</sub>).
- Calculating subpixel spectral diversity from  $s_i$  and  $\alpha_i$  using the • spectral species richness and Simpson index.
- Using subpixel spectral diversity as a proxy of plant community diversity.

### **Implications:**

- Results obtained from DESIS data have been encouraging, indicating the potential of forthcoming spaceborne imagers to map plant diversity.
- Further developments and tests on different ecosystems and datasets are needed to operationalize the approach.

### Capturing plant diversity with spaceborne imaging spectroscopy



DESIS data at 30m resolution

Spectral endmember abundance





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Endmember diversity

## **Results from simulated data:** 15,000 simulated

communities (3-17 prairie grassland species per community and soil; signal-to-noise ratio of 60).



**Results from real-world DESIS data:** Significant relationship between endmember diversity from DESIS data and *in-situ* measured taxonomic Simpson index and phylogenetic evenness in prairie grasslands. These results are from 100 240 m × 240 m plots and their corresponding spectra at the Joseph H. Williams Tallgrass Prairie Preserve, Oklahoma.





Endmember Simpson index